

→ Now that we have Green's fn we can get back to the antenna problem. (lect 05)

Recall: $\nabla^2 \vec{A} + k^2 \vec{A} = -\mu \vec{J}$ and $\vec{J} = J_z \hat{z}$

∴ we wanted to solve $\nabla^2 A_z + k^2 A_z = -\mu J_z(r)$

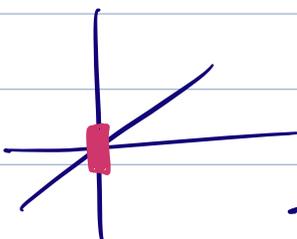
We now know $A_z(r) = \int_{-\infty}^{\infty} d^3 r' g(r, r') (\mu J_z(r'))$

Similarly when solving $\nabla^2 \vec{F} + k^2 \vec{F} = -\epsilon \vec{M}$

we will get $\vec{F}(r) = \int_{-\infty}^{\infty} d^3 r' g(r, r') (\epsilon \vec{M}(r'))$

↳ Now let's study an imp special case, the Hertz dipole, a.k.a simplest antenna. [Insights from here extend to all antennas]

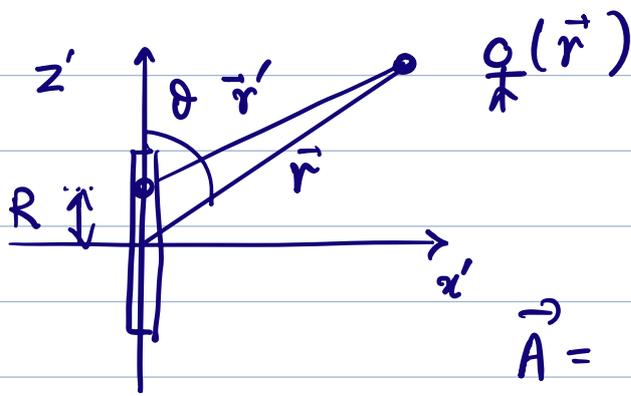
Assume: $J_z(r) = \begin{cases} I_0 \delta(x) \delta(y) \hat{z}, & |z| < l \\ 0, & \text{else} \end{cases}$



further assume $l \ll \lambda$. $\left(\frac{2\pi}{\lambda}\right)$

Soln: $A_z(r) = \int_{-l/2}^{l/2} \frac{I_0 \mu e^{-jk|r'-r|}}{4\pi|r'-r|} dz'$

Recall the geometry for this integral.



Since $|r'| \ll \lambda$
This becomes:

$$\vec{A} = \frac{I_0 \mu_0 l}{4\pi r} e^{-jkr} \hat{z}$$

Makes more sense to just put it into (r, θ, ϕ)

$$\therefore A_r = A_z \cos \theta = \frac{\mu_0 I_0 l}{4\pi r} e^{-jkr} \cos \theta$$

$$A_\theta = -A_z \sin \theta = -\frac{\mu_0 I_0 l}{4\pi r} \sin \theta e^{-jkr}$$



$$A_\phi = 0$$

Now let's get the fields: $\vec{H} = \frac{1}{\mu} \nabla \times \vec{A}$

$$= \frac{1}{\mu} \hat{\phi} \left[\frac{1}{r} \left(\frac{\partial}{\partial r} (r A_\theta) - \frac{\partial A_r}{\partial \theta} \right) \right]$$

$$\text{Substituting we get, } H_\phi = j \frac{k I_0 l \sin \theta}{4\pi r} \left(1 + \frac{1}{jkr} \right) e^{-jkr}$$

$$\text{Remember } \vec{E} = \frac{1}{j\omega \epsilon} \nabla \times \vec{H}$$

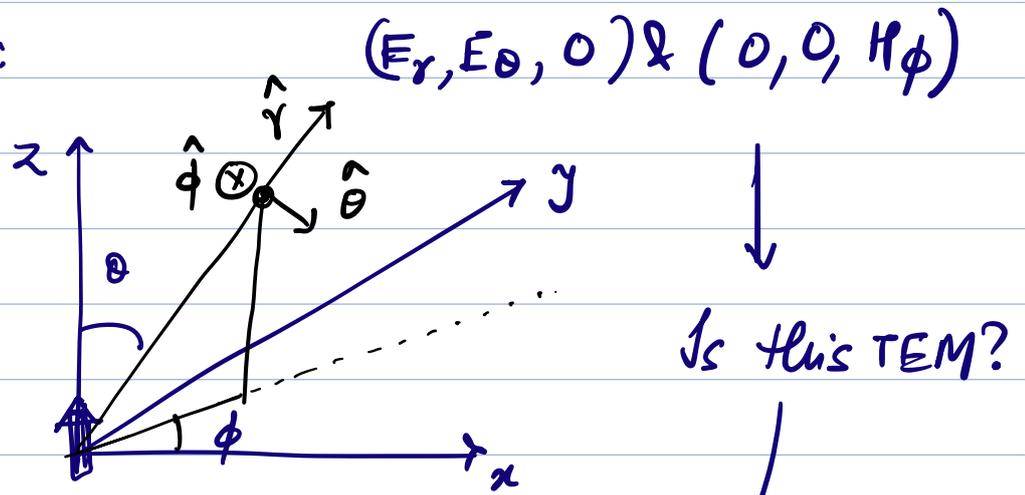
Using $\nabla \times$ in spherical coordinates.

$$E_r = \eta \frac{I_0 l \cos \theta}{2\pi r^2} \left(1 + \frac{1}{jkr} \right) e^{-jkr}$$

$$E_\theta = j \eta \frac{k I_0 l \sin \theta}{4\pi r} \left(1 + \frac{1}{jkr} - \frac{1}{(kr)^2} \right) e^{-jkr}$$

$$E_\phi = 0.$$

So, we have:



Can't tell right away, need to see where the Poynting vector is!

$$S = \frac{1}{2} E \times H^* \quad (\text{complex Poynting vector})$$

$$= \frac{1}{2} (\hat{r} E_r + \hat{\theta} E_\theta) \times (\hat{\phi} H_\phi^*)$$

$$S = \frac{1}{2} (\hat{r} E_\theta H_\phi^* - \hat{\theta} E_r H_\phi^*)$$

Substituting: $S_r = \frac{\eta}{8} \left| \frac{I_{0l}}{\lambda} \right|^2 \frac{\sin^2 \theta}{r^2} \left(1 - j \frac{1}{(kr)^3} \right)$

$$S_\theta = j\eta \frac{k |I_{0l}|^2 \cos \theta \sin \theta}{16\pi r^3} \left(1 + \frac{1}{(kr)^2} \right)$$

Interesting what happens when we integrate this power over an arbitrary sphere.

$$P = \iint \vec{S} \cdot d\vec{s} = \int_0^{2\pi} \int_0^\pi (\hat{r} S_r + \hat{\theta} S_\theta) \cdot \hat{r} r^2 \sin \theta d\theta d\phi$$

↑
nothing

Gives: $P = \int_0^{2\pi} \int_0^{\pi} S_r r^2 \sin\theta d\theta d\phi$

No approx here about r being small or large!

Time avg \rightarrow $= \hat{r} \frac{\eta \pi}{3} \left| \frac{I_0 l}{\lambda} \right|^2 \left[1 - j \frac{1}{(kr)^3} \right]$

Power \rightarrow $P_{rad} = \text{Re}[P]$
 $= \frac{\eta \pi}{3} \left| \frac{I_0 l}{\lambda} \right|^2 = \frac{1}{2} |I_0|^2 R_r$

R_r is an equivalent qty, called the radiation resistance of the antenna.

↳ For this dipole antenna, $R_r \approx 80\pi^2 \left(\frac{l}{\lambda}\right)^2$

if $l = \frac{\lambda}{50}$, $R_r \approx 0.32 \Omega$

Terrible antenna, since far from 50Ω matching line! HW: what is Γ ?

↳ Let's get back to the fields.

Field terms vary as $\frac{1}{r}$, $\frac{1}{r^2}$, $\frac{1}{r^3}$

far field or radiation field
 (only term that survives far away)

Inductive field
 (Biot Savart law)

electrostatics field,
 (dipole charge like eqn.)

[Near field $kr \ll 1$]

[Far field $kr \gg 1$]

$$E_r \propto j/r^3$$

0 (relatively)

$$E_\theta \propto j/r^3$$

$$j/r$$

$$H_\phi \propto 1/r^2$$

$$j/r$$

$$S \propto \frac{1}{r^2}, \frac{j}{r^3}, j/r^5$$

$$1/r^2$$

$$S_{av} \propto \frac{1}{r^2} \hat{r} / 0$$

$$\frac{1}{r^2} \hat{r}$$

OR
(depending on your view)

\Rightarrow TEM wave in farfield.

Rough approximate boundary is $r_0 = \frac{2D^2}{\lambda}$

where D is max dimension of the antenna.

— x —

We saw a specific case of an antenna derivation (dipole with specific current distrib).

So, in general for a \vec{J} in a bounded region, what all can we say?

1) From $\vec{J} \rightarrow$ we get \vec{A} using g .

2) form of \vec{A} ? $\propto \frac{1}{r^n}$, $n=1$ far field (Saw for Hertz)
 $n>1$ near field

$$\text{i.e. } \vec{A} \approx \left[\hat{r} A'_r(\theta, \phi) + \hat{\theta} A'_\theta(\theta, \phi) + \hat{\phi} A'_\phi(\theta, \phi) \right] \frac{e^{-jk r}}{r}$$

Most antenna problems, we are interested in the far zone behaviour. ← (for large r)

Like we saw in the dipole case, $A'_r(\theta, \phi) \approx 0$. $\left[\text{Has } \frac{1}{r^n}, n > 1 \right]$

$$\text{Subst. } \vec{A} \text{ into } \vec{E} = -j\omega \vec{A} - \frac{j}{\omega\mu\epsilon} \nabla(\nabla \cdot \vec{A})$$

Again, under far field approx, this turns out to be:

$$\vec{E}_A = -j\omega \vec{A} \Rightarrow E_r = 0$$

$$E_\theta \approx -j\omega A_\theta$$

$$E_\phi \approx -j\omega A_\phi$$

$$H_r \approx 0$$

$$H_\theta \approx \frac{j\omega A_\phi}{\eta} = -\frac{E_\phi}{\eta}$$

$$H_\phi \approx \frac{-j\omega A_\theta}{\eta} = \frac{E_\theta}{\eta}$$

$$H_A \approx \frac{\hat{r} \times \vec{E}_A}{\eta} \Rightarrow \text{TEM wave}$$

[Q: How many poles hidden here?]

Similarly we can write for $H_F = -j\omega \vec{F}$ & so on.

The above expressions are powerful as they apply to all bounded source (in space) antennas in the far field.